

*Special INNS Workshop: International Neural Network Society
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Morphological Classification of Radio Galaxies Using Semi-Supervised Group Equivariant CNNs

Authors

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Introduction

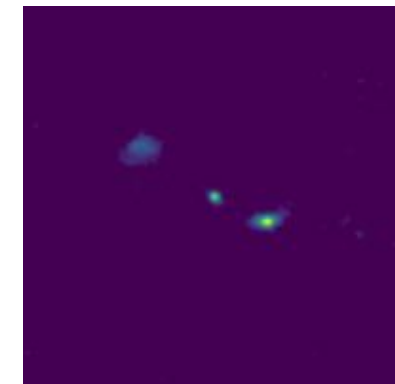
- The cosmos teems with a diverse array of galaxies, providing rich data for understanding the universe.
- Galaxy classification is crucial for understanding the formations and evolution of galaxies.
- Manual classification of galaxies is becoming impractical due to the sheer number of detected galaxies.
- The necessity of machine learning arises to handle the vast amount of data and automate the classification process.
- Current challenges include **handling the diverse orientations of galaxies** and **efficiently using large amounts of unlabeled data**.

Radio Galaxies: Types and Morphology

- Radio galaxies emit significant radiation at radio wavelengths.
- Two main types of radio galaxies based on morphology: Fanaroff-Riley Type I (FRI) and Type II (FR II).
- FRI radio galaxies have a large, **bright core and diffuse, extended lobes**. The lobes are elongated, connected to the core, and exhibit a fainter radio emission away from the core.
- FR II radio galaxies have **brightened lobes and a fainter core**. They are typically more luminous than FRI radio galaxies, with observed hotspots at the ends of the lobes.



FRI Galaxy



FR II Galaxy

Current Methods for Galaxy Classification and Drawbacks

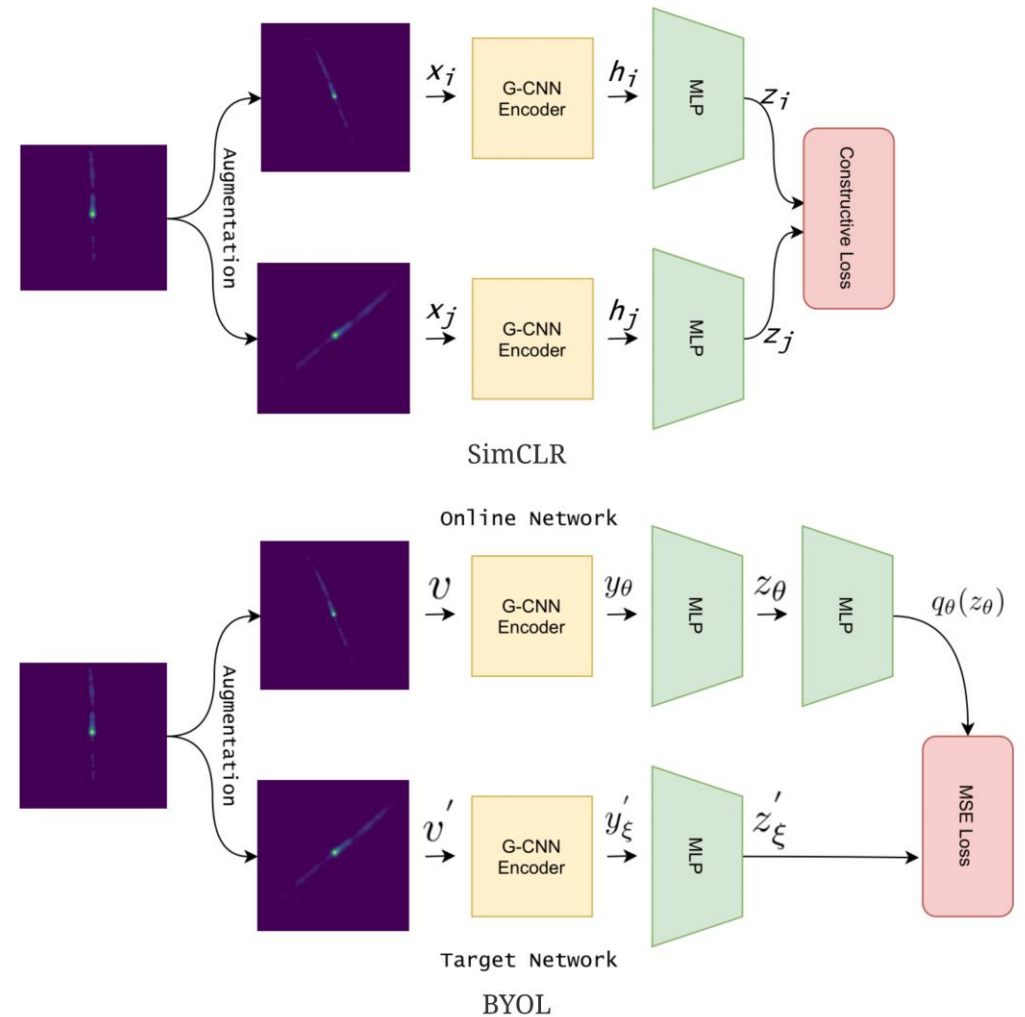
- Various machine-learning techniques, including CNNs and Region-based CNNs, are used for classifying radio galaxies.
- Challenges include handling different galaxy orientations and isometries like translation, rotation, and mirror reflection.
- Augmenting data with rotated images has limitations in improving classification accuracy.
- Effective utilization of a large amount of unlabeled galaxy data remains a challenge in the classification process.
- Semi-supervised learning holds promise in extracting meaningful information from unlabeled data, but further advancements are needed to fully exploit its potential

Objective of the study

- Develop an effective classification approach for radio galaxies using a semi-supervised learning framework.
- Utilize self-supervised learning techniques to learn robust representations from unlabeled data.
- Fine-tune the learned representations using a labeled dataset for accurate galaxy classification.
- Address the challenge of diverse galaxy orientations using Group Equivariant Convolutional Neural Networks (G-CNNs).

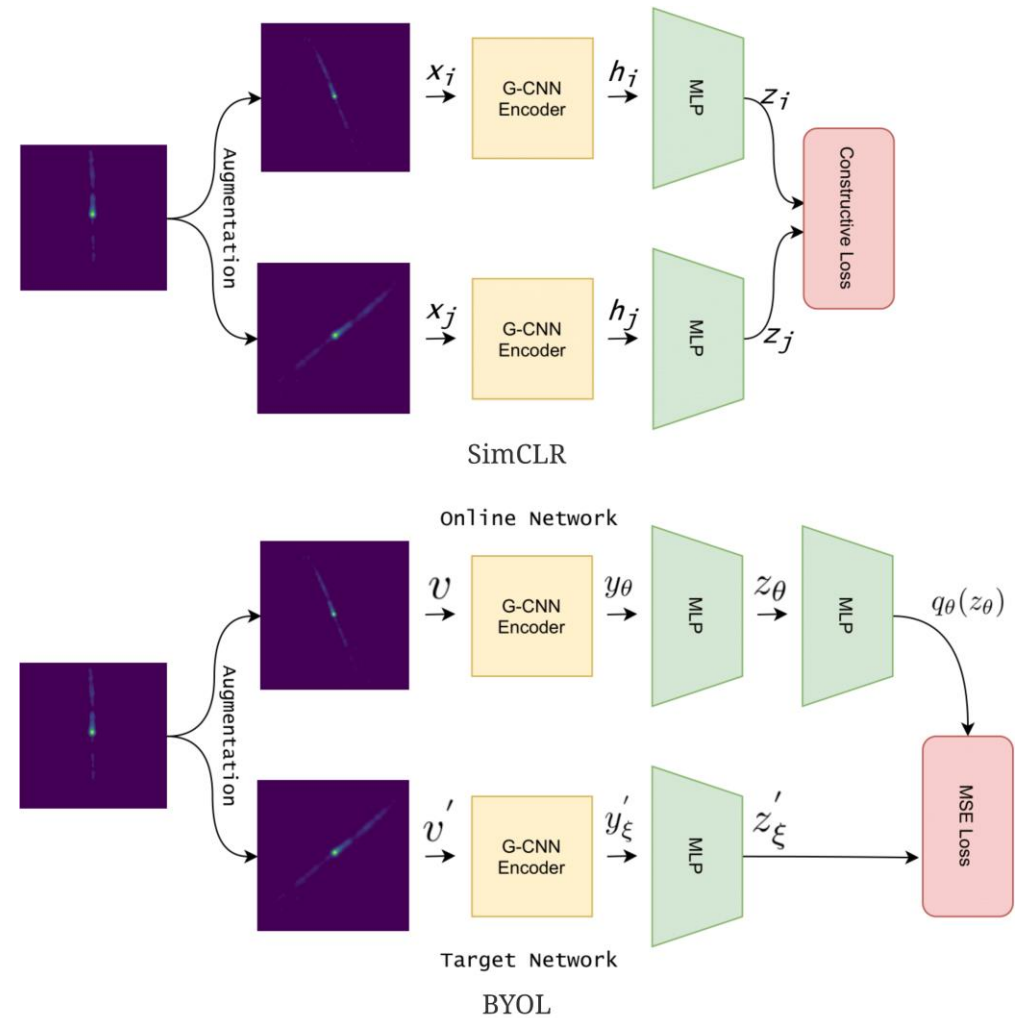
Proposed Method

- Utilize self-supervised learning techniques, such as SimCLR and BYOL, to learn representations from a large unlabeled dataset.
- Extract meaningful features from the data through self-supervised learning.
- Apply contrastive learning frameworks to enhance the model's ability to capture important patterns and structures.



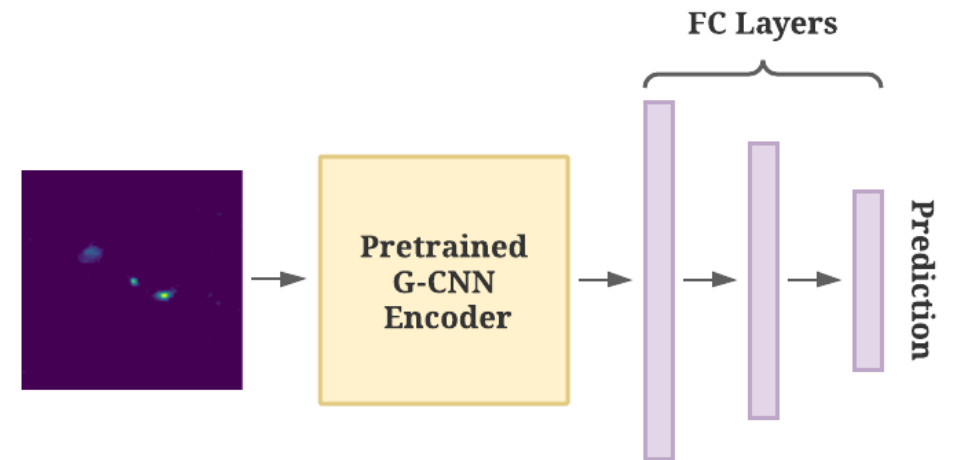
Proposed Method

- Modify the encoders of self-supervised models to handle diverse galaxy orientations.
- Employ Group Equivariant Convolutional Neural Networks (G-CNNs) as feature extractors.
- Ensure the model remains invariant to different isometries, such as translation, rotation, and mirror reflection.



Proposed Method

- Modify the pre-trained encoder obtained from self-supervised learning.
- Replace the last fully connected layer of the encoder with a new architecture of three fully connected linear layers.
- Adapt the model to the specific task of FR classification.
- Employ a supervised learning approach with cross-entropy loss for fine-tuning.



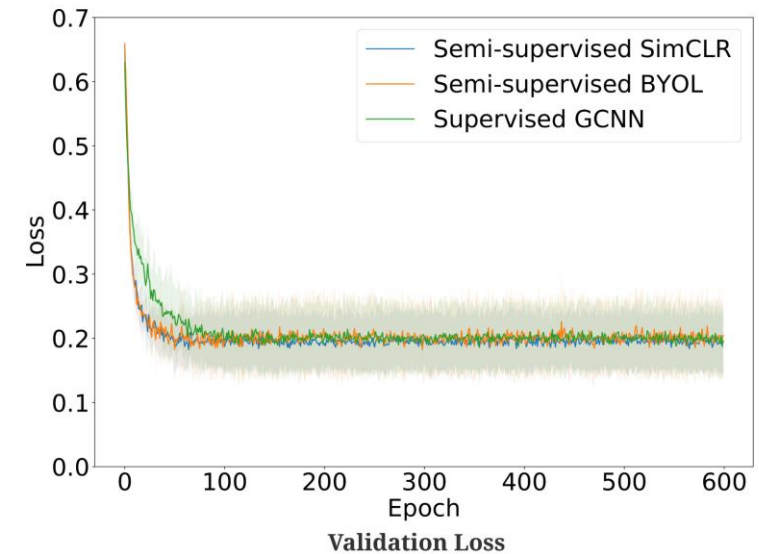
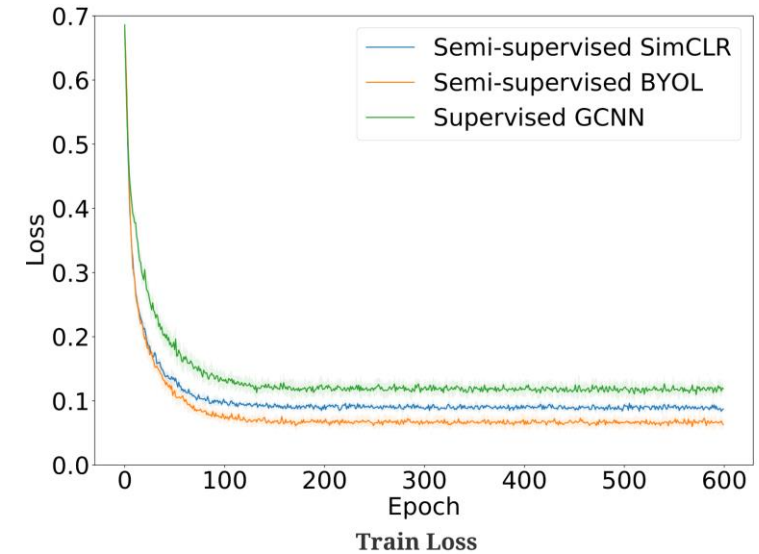
Results - Cluster Quality Analysis

- Visualizations of representations obtained from fine-tuned encoders show distinct clustering of FRI and FR II classes
- Semi-supervised models (BYOL and SimCLR) exhibit significantly improved cluster quality compared to the supervised model
- Silhouette Score and Davies Bouldin Score confirm the effectiveness of the semi-supervised models in capturing meaningful patterns and structures



Results - Convergence Analysis

- Evaluated convergence during fine-tuning using 5-fold cross-validation on Dataset-F.
- Rapid convergence observed in training and validation loss plots.
- Fine-tuned encoders converged faster than the supervised G-CNN approach.
- Demonstrates efficiency and effectiveness of our semi-supervised learning approach for FR classification.



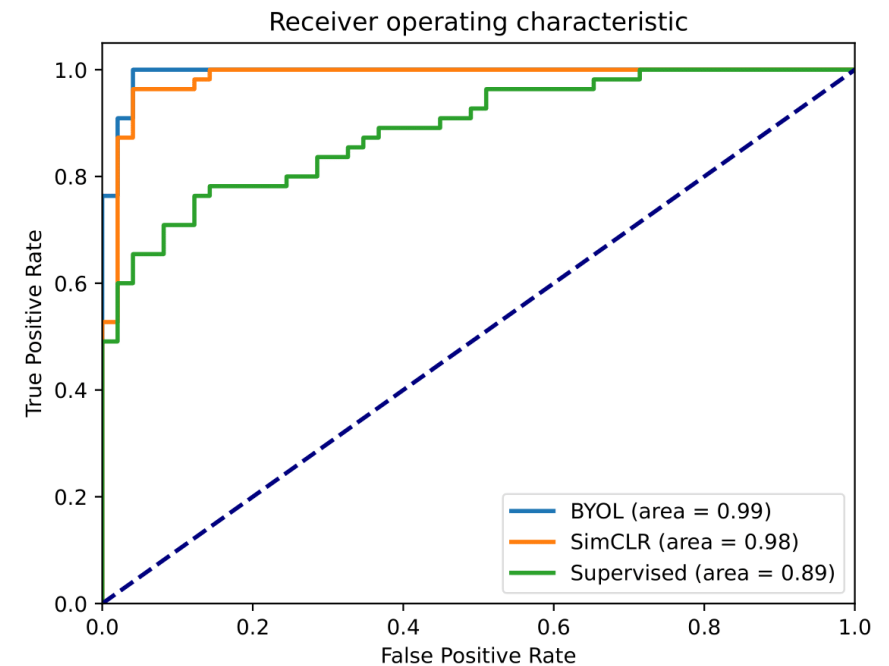
Results - Performance Comparison

Table 3. Performance comparison between Semi-supervised and Supervised methods; The table illustrates the superiority of our Semi-supervised models over the state-of-the-art supervised method across various classification metrics

	Accuracy[%]	FRI			FRII		
		Precision	Recall	f1-score	Precision	Recall	f1-score
Semi-supervised SimCLR	<u>95.77 ± 0.90</u>	0.98 ± 0.061	0.93 ± 0.018	<u>0.95 ± 0.011</u>	0.94 ± 0.013	0.98 ± 0.014	<u>0.96 ± 0.009</u>
Semi-supervised BYOL	97.12 ± 0.40	<u>0.97 ± 0.008</u>	<u>0.96 ± 0.009</u>	0.97 ± 0.005	<u>0.96 ± 0.007</u>	<u>0.98 ± 0.008</u>	0.97 ± 0.004
Supervised G-CNN	94.80 ± 0.90	0.93 ± 0.012	0.96 ± 0.010	0.94 ± 0.009	0.96 ± 0.009	0.94 ± 0.012	0.95 ± 0.009

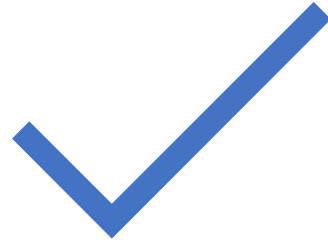
Results - ROC Analysis and Statistical Significance

- Utilized ROC curves and AUC scores for performance evaluation.
- AUC scores of 0.99 (BYOL) and 0.98 (SimCLR) indicate accurate distinction between radio galaxy classes.
- Minimal false positives ensure correct identification of radio galaxies.
- Statistical t-test shows significant improvement in performance compared to supervised G-CNN.
- t-test for semi-supervised BYOL: t-value of approximately -3.47, p-value of approximately 0.0038.



Conclusion

- Our novel approach for classifying radio galaxies demonstrated significant effectiveness.
- Fine-tuning the pre-trained encoder with limited labeled data improved performance over traditional supervised methods.
- Our semi-supervised models outperformed the supervised model in accuracy, convergence, and ability to distinguish between classes.
- Our work highlights the potential of semi-supervised learning in radio galaxy classification and encourages further research in the field.



Thank You